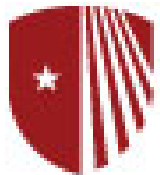


PHY 452/562
Lasers and Modern Optics

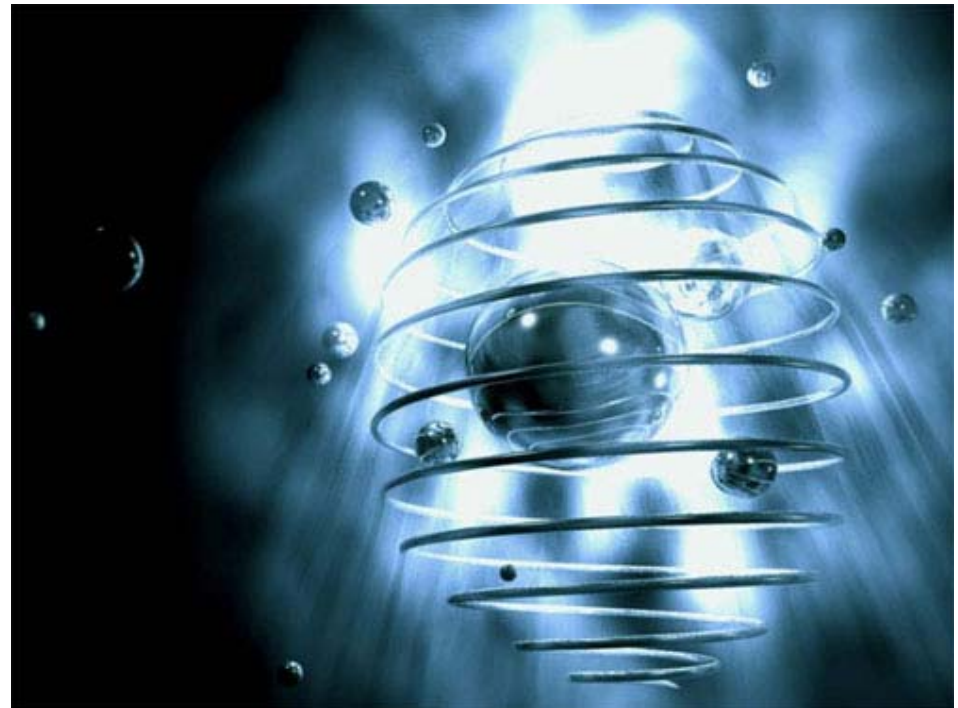
Prof. Eden Figueroa

Lecture 6: Quantum Mechanics I

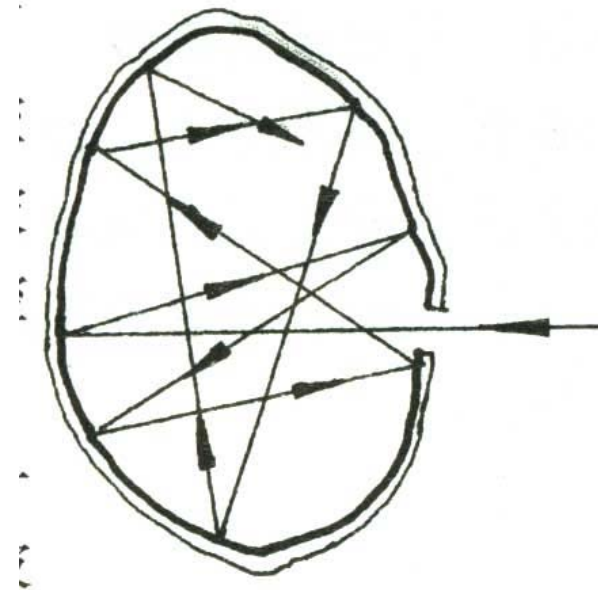
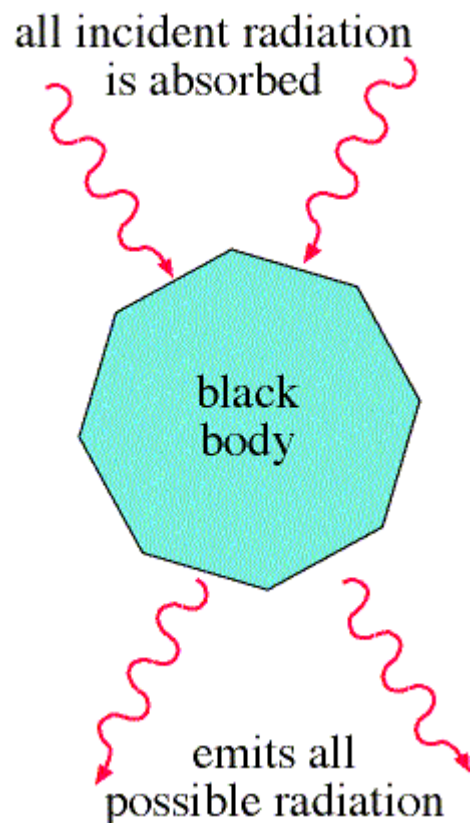


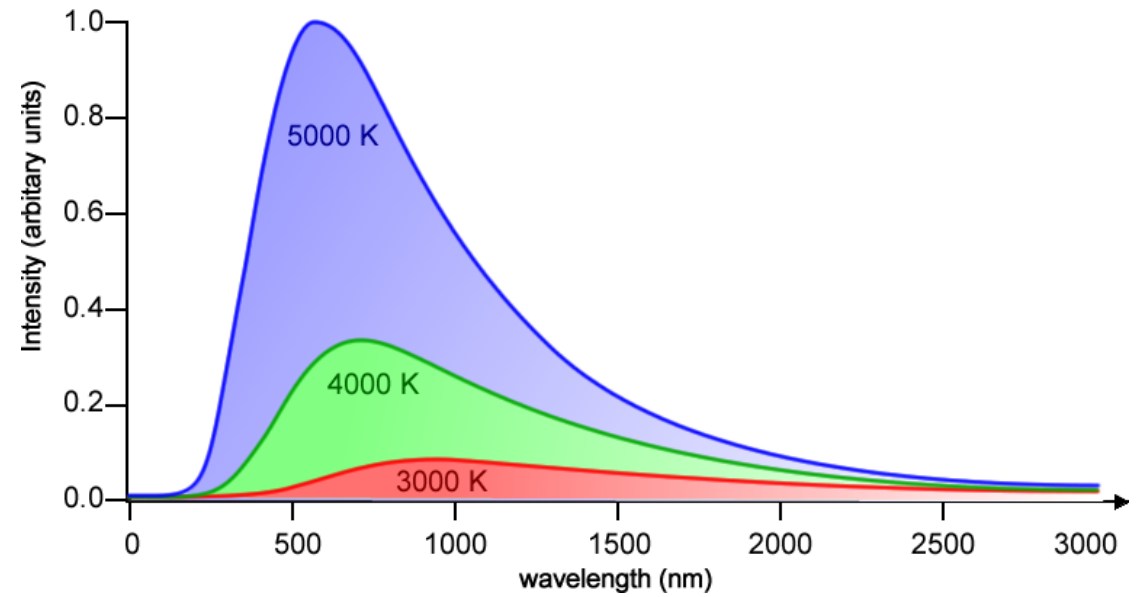
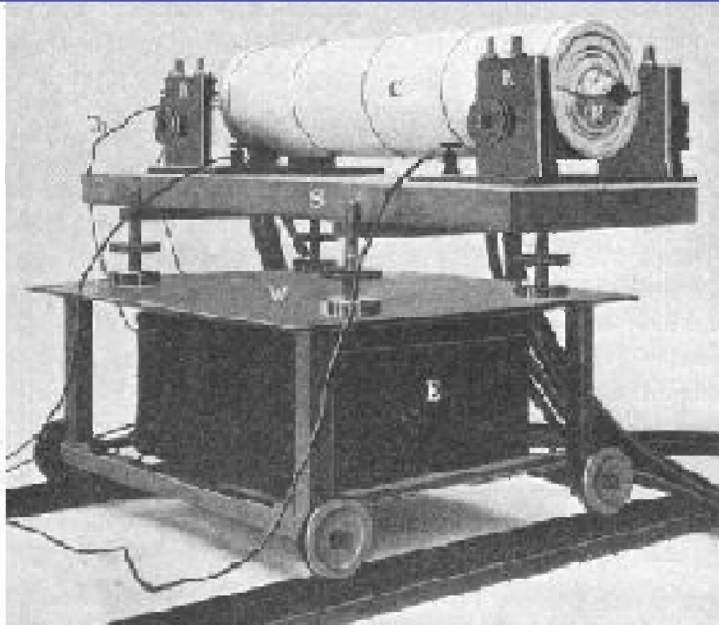
Stony Brook
University

- Black body radiation.
- Quantization a la Planck.



- Blackbody is defined as the body which can absorb all energies. A large cavity with a small hole on its wall can be taken as a blackbody.





Lummer and Kurlbaum's black-body experiment from 1898



Stefan and Boltzmann's law: the radiation energy is proportional to the fourth power of the associated temperature.

Stefan-Boltzmann Law

Power radiated
(Watts)

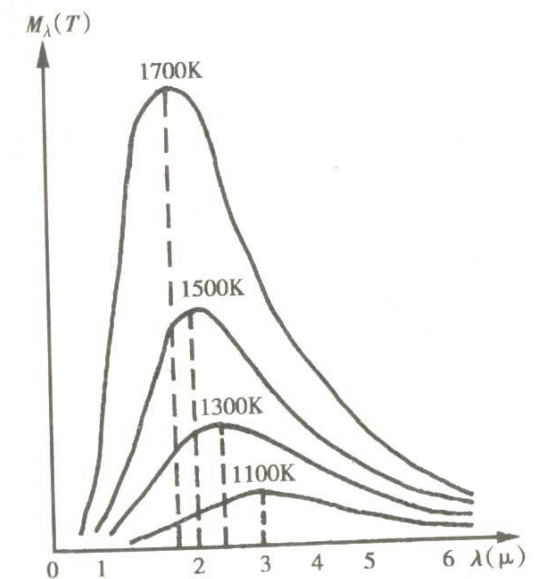
emissivity
(no units)

Surface area
(m²)

$$P = e\sigma AT^4$$

Stefan-Boltzmann constant
5.67x10⁻⁸ W m⁻² K⁻⁴

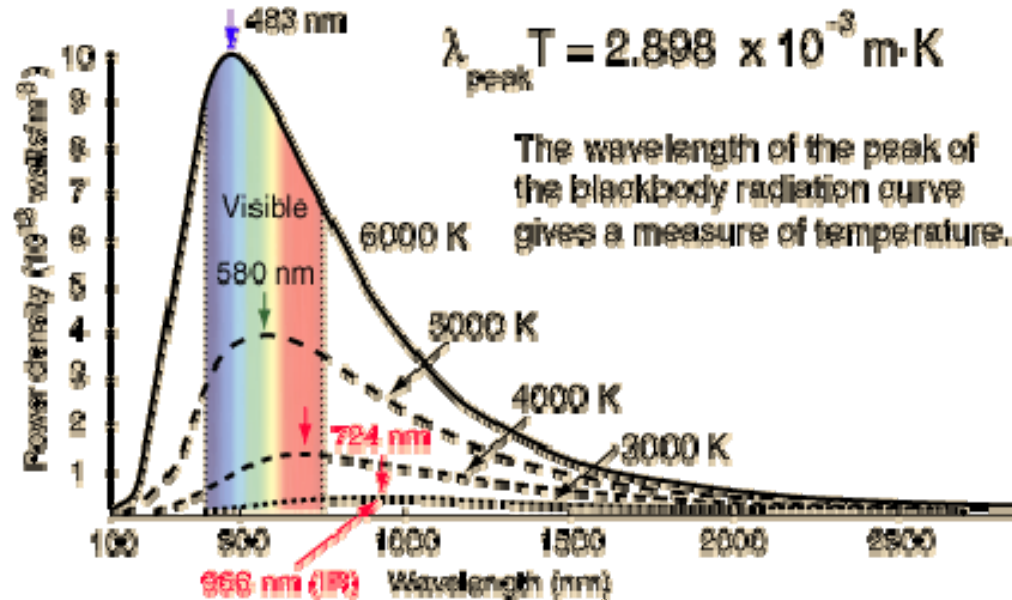
Temperature
(Kelvins)



Blackbody radiation of spectra for four different temperatures.

Wien's displacement law: the peak of the curve shifts towards longer wavelength as the temperature falls and it satisfies

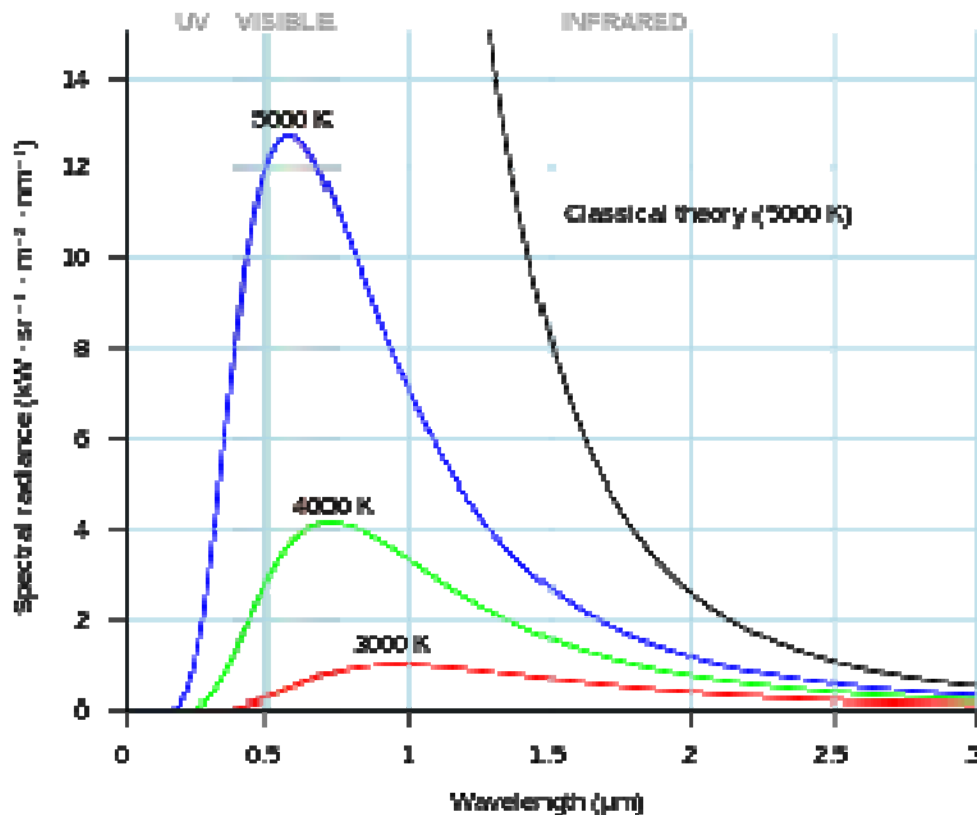
$$\lambda_{peak} T = b \quad \text{where } b \text{ is called the Wein's constant.}$$



This law is quite useful for measuring the temperature of a blackbody with a very high temperature.

Rayleigh and Jeans theory

In 1890, Rayleigh and Jeans obtained a formula using the classical electromagnetic (Maxwell) theory and the classical equipartition theorem of energy:

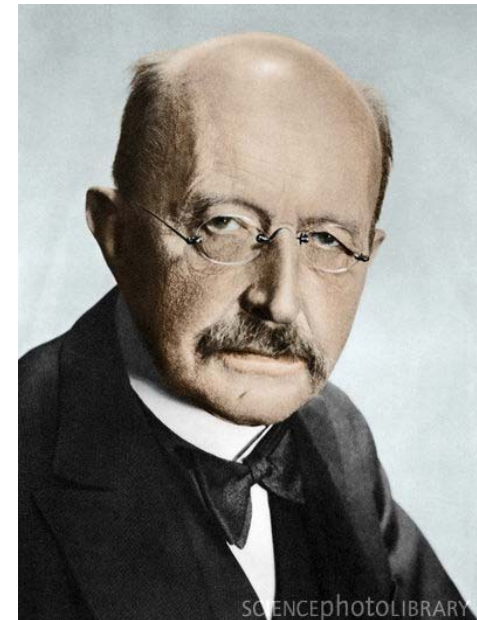


$$M_{\lambda}(T) = C_1 \lambda^{-4} T$$

Planck's formula

In 1900, after studying the above two formulas carefully, Planck proposed an empirical formula.

$$M_{\lambda}(T) = 2\pi hc^2 \lambda^{-5} \frac{1}{e^{\frac{hc}{k\lambda T}} - 1}$$



Where c is the speed of light, k is Boltzmann's constant, h is Planck's constant and e is the base of natural logarithm.

The formula can describe the curve of blackbody radiation exactly for all wavelengths.

- For very large wavelength, the Rayleigh-Jeans formula can be obtained from Planck's formula;

$$\frac{hc}{k\lambda T} \ll 1$$

$$\frac{hc}{k\lambda T} \ll 1$$

$$e^{\frac{hc}{k\lambda T}} = 1 + \frac{hc}{k\lambda T} + \frac{1}{2} \left(\frac{hc}{k\lambda T} \right)^2 + \dots$$

Drop the second order and higher order terms, and RJ formula could be obtained.

For smaller wavelength of blackbody radiation, the Wein's formula can be achieved also from Planck's formula;

$$\frac{1}{e^{\frac{hc}{k\lambda T}} - 1} \approx e^{-\frac{hc}{k\lambda T}}$$

Integrating Planck's formula with respect to wavelength, the Stefan and Boltzmann's law can be obtained as well.

$$M(T) = \int_0^{\infty} M_{\lambda}(T) d\lambda = \sigma T^4$$

Planck's Hypotheses:

- The molecules and atoms composing the blackbody can be regarded as the linear harmonic oscillator with electrical charge;
- **The oscillators can only be in a special energy state. All these energies must be the integer multiples of a smallest energy ($\epsilon_0 = h\nu$). Therefore the energies of the oscillators are $E = n h\nu$ with $n = 1, 2, 3, \dots$**